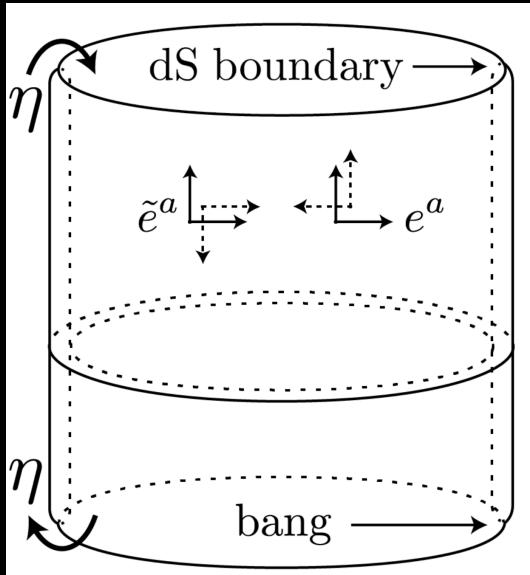
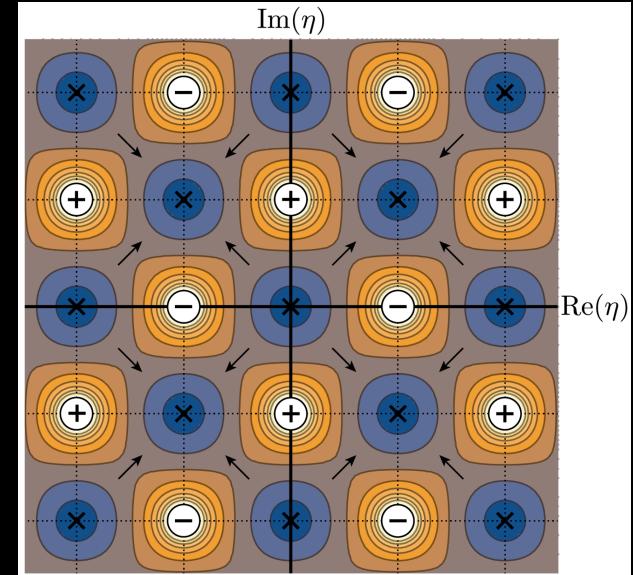


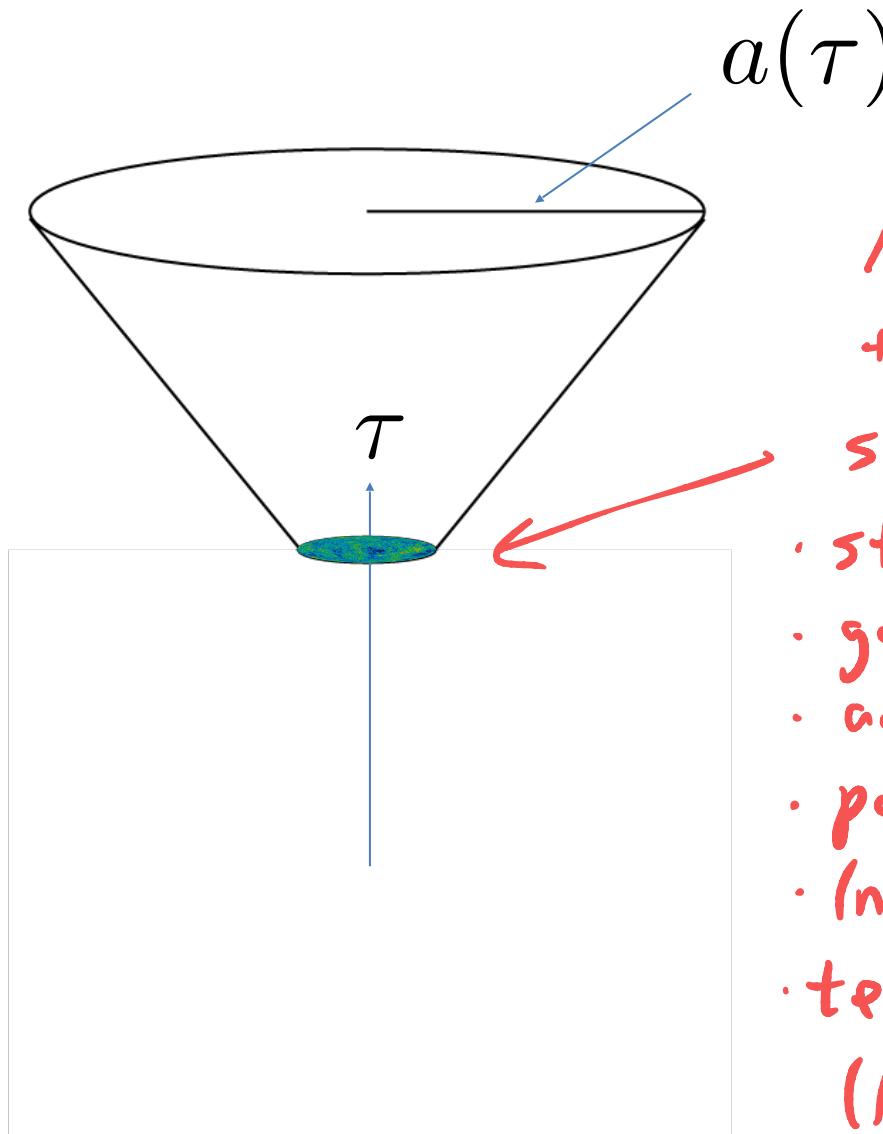
# What is the simplicity of the early universe trying to tell us?



Latham Boyle  
Perimeter Institute

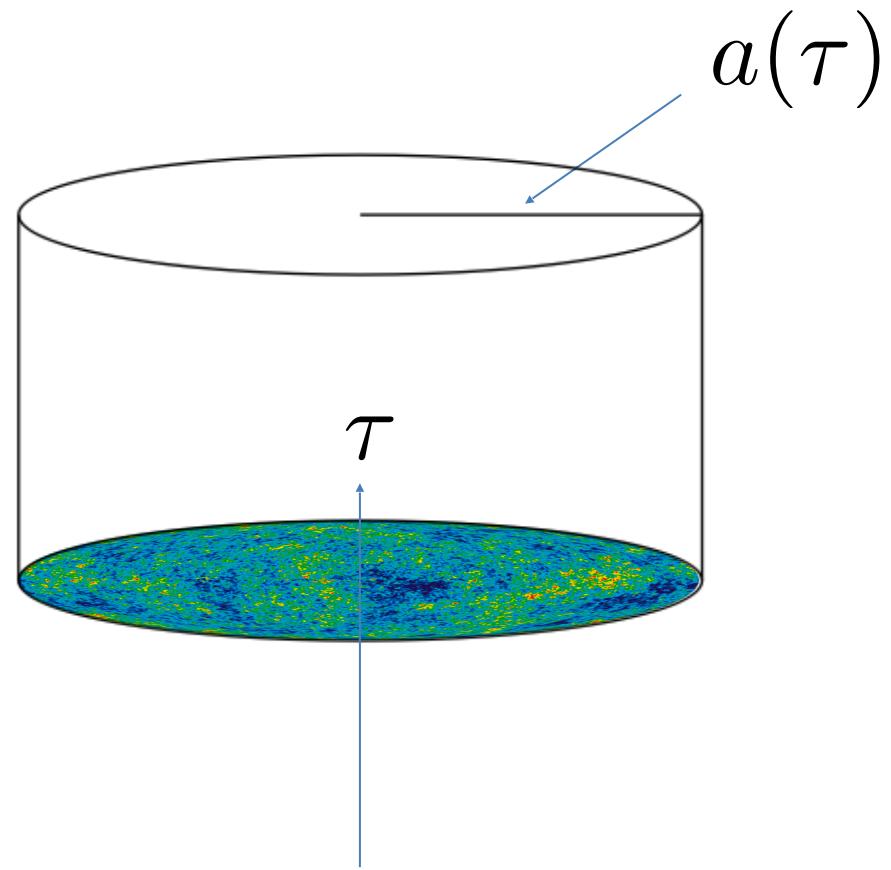


- 1) LB, K. Finn and N. Turok, *CPT-Symmetric Universe* arXiv:1803.08928 (Phys. Rev. Lett.)
- 2) LB, K. Finn and N. Turok, *The Big Bang, CPT, and neutrino dark matter* arXiv:1803.08930 (Ann. Phys.)
- 3) LB and N. Turok, *Two-Sheeted Universe, Analyticity & Arrow of Time*, arXiv:2109.06204
- 4) LB and N. Turok, *Cancelling the Vacuum Energy and Weyl Anomaly in the Standard Model with Dimension-Zero Scalar Fields*, arXiv:2110.06258
- 5) N. Turok and LB, *Gravitational entropy and the flatness, homogeneity and isotropy puzzles*, arXiv:2201.07279
- 6) LB and N. Turok, *Thermodynamic solution to the homogeneity, isotropy and flatness puzzles (and a clue to the cosmological constant)*, arXiv:2210.01142
- 7) N. Turok and LB, *A Minimal Explanation of the Primordial Cosmological Perturbations*, arXiv:2302.00344

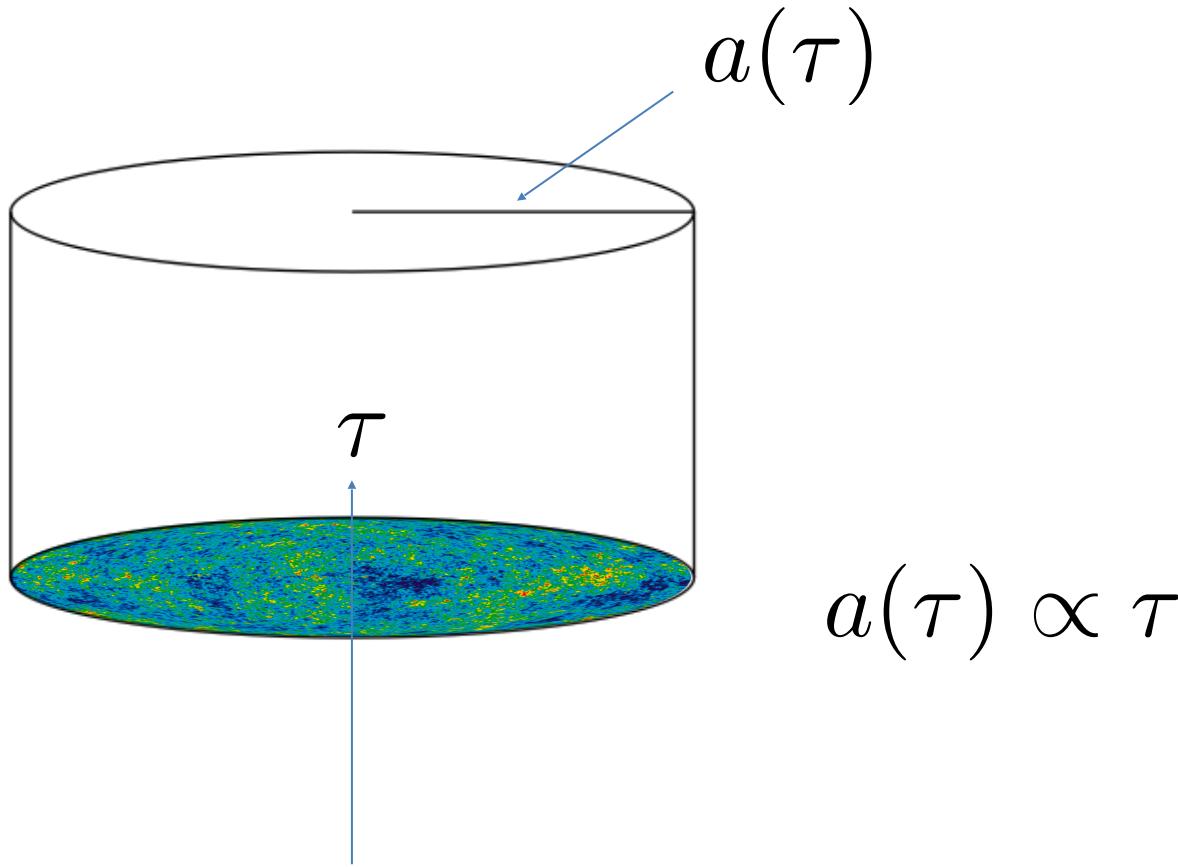


Maximally symmetric  
+ tiny ( $10^{-5}$ ) random  
scalar (density) perturbations:

- statistically symmetric,
- gaussian,
- adiabatic,
- power law,
- (nearly) scale invariant,
- temporally synchronized  
(Neumann b.c.'s for  $\zeta$   
at the bang)

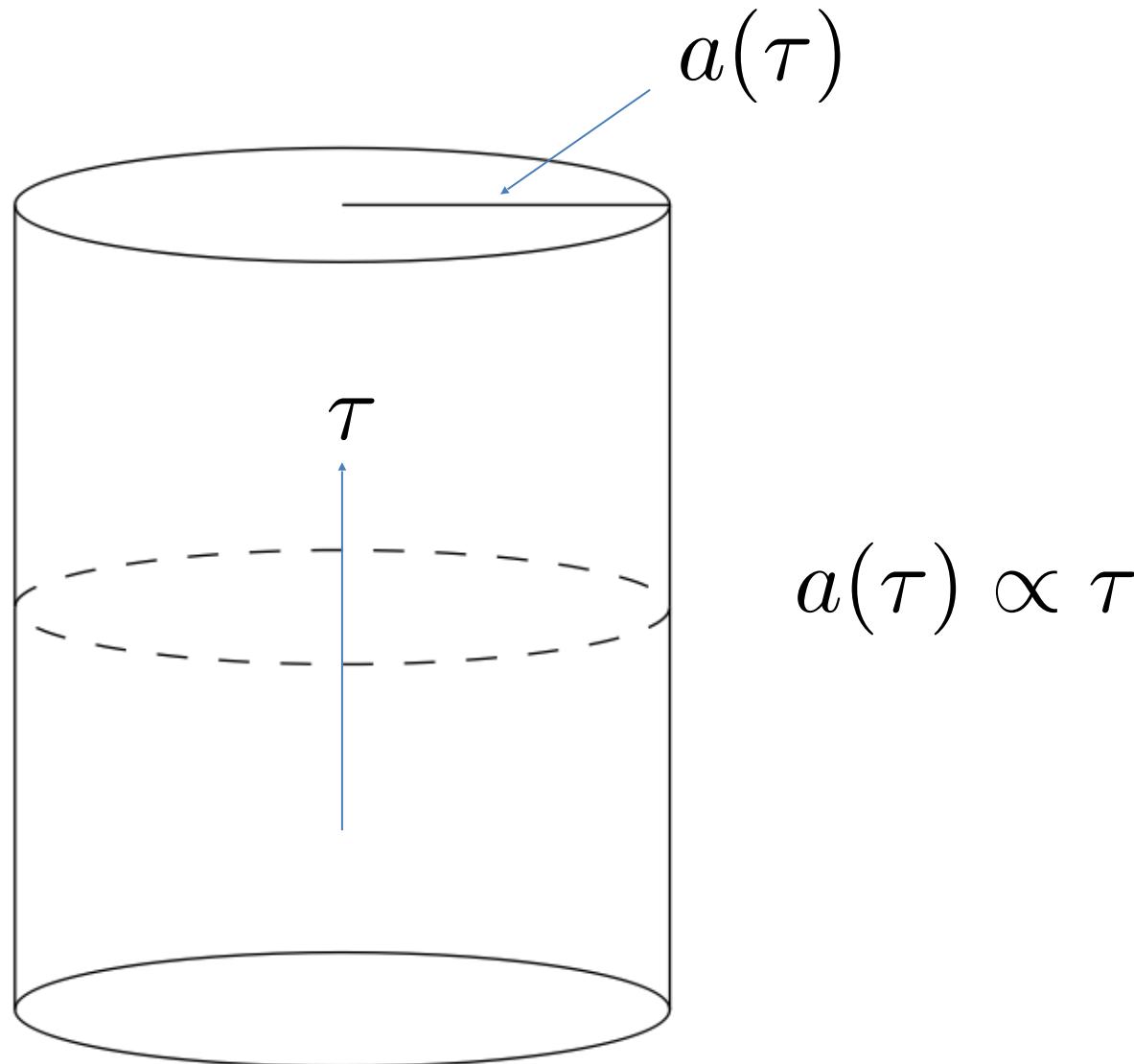


$$g_{\mu\nu} = a^2(\tau) \eta_{\mu\nu}$$



$$a(\tau) \propto \tau$$

$$g_{\mu\nu} = a^2(\tau) \eta_{\mu\nu}$$

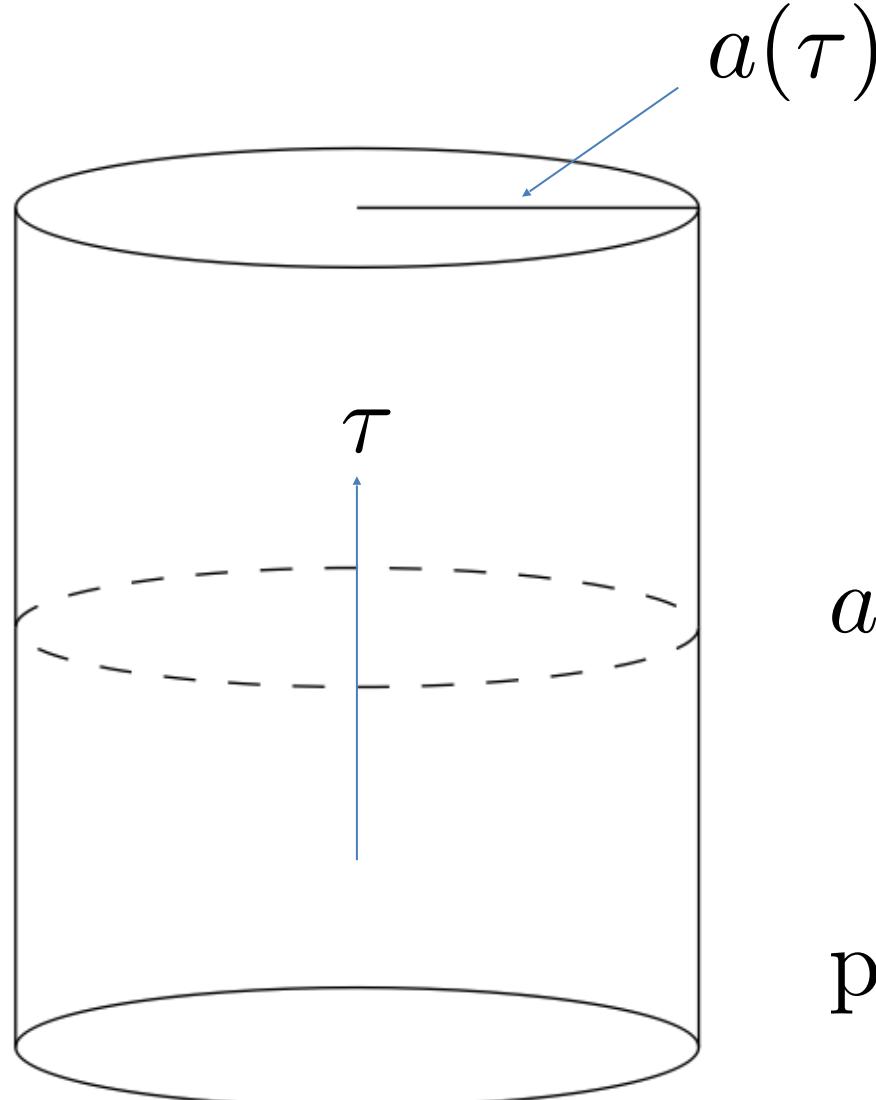


$$a(\tau) \propto \tau$$

$$g_{\mu\nu} = a^2(\tau) \eta_{\mu\nu}$$

new isometry:

$$\tau \rightarrow -\tau$$



$$a(\tau) \propto \tau$$

preferred vacuum:

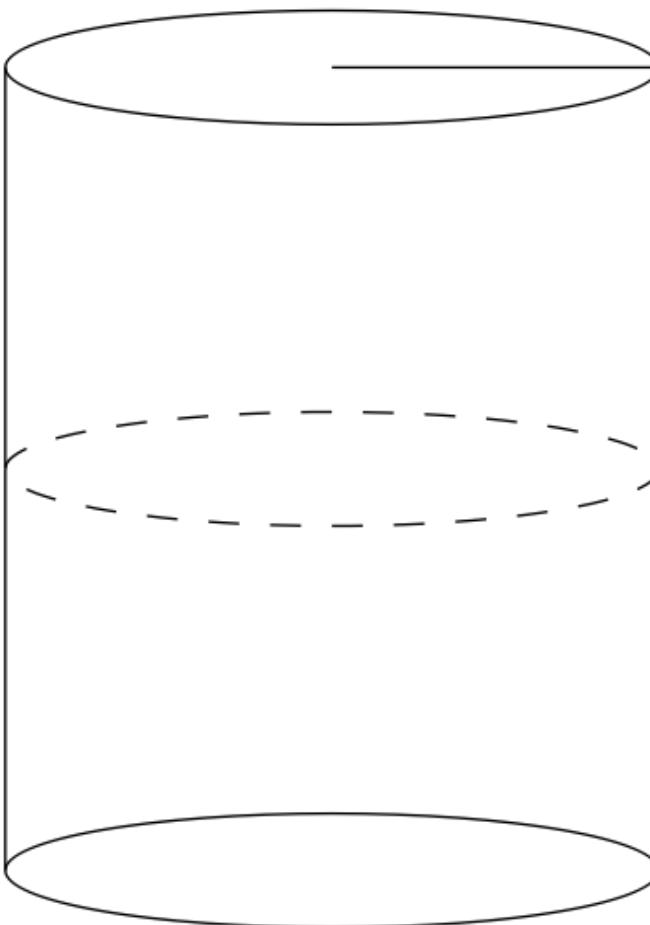
$$|0_{CPT}\rangle$$

hypothesis:

the universe does not spontaneously violate CPT

$$\psi(x) = \sum_h \int \frac{d^3\mathbf{p}}{(2\pi)^{3/2}} [a(\mathbf{p}, h)\psi(\mathbf{p}, h, x) + b^\dagger(\mathbf{p}, h)\psi^c(\mathbf{p}, h, x)]$$

$$\psi_+(\mathbf{p},h,x)$$



$$a_+, b_+ \Rightarrow |0_+\rangle$$

$$\psi_0(\mathbf{p},h,x)$$

$$\left( \, (\psi_0(\tau) \sim \psi_0^c(-\tau) \, \right)$$

$$\psi_-(\mathbf{p},h,x)$$

$$a_0, b_0 \Rightarrow |0_0\rangle$$

$$a_-, b_- \Rightarrow |0_-\rangle$$

# the standard model

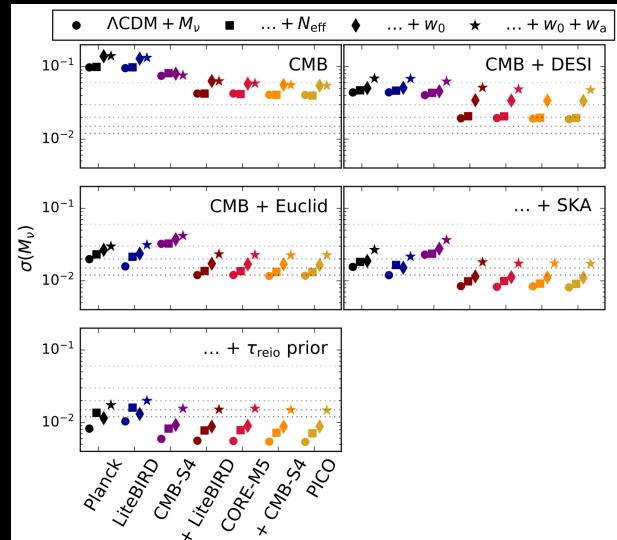
$$G_\mu, W_\mu, B_\mu, h$$

$$\left. \begin{array}{c} d_L, u_L, d_R, u_R \\ d_L, u_L, d_R, u_R \\ d_L, u_L, d_R, u_R \\ e_L, \nu_L, e_R, \nu_R \end{array} \right\} \times 3$$

Prediction 0: dark matter neutrino is  $4.8 \times 10^8$  GeV.

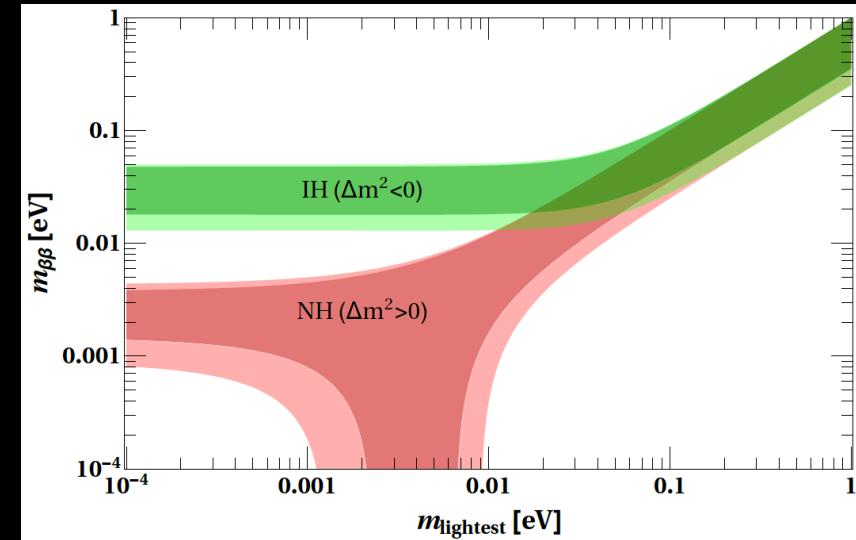
Prediction 1: one neutrino is massless

$$\sum m_{\nu} \approx .06 \text{ eV (NH)} \text{ or } .12 \text{ eV (IH)}$$



(Brinckmann et al, arXiv:1808.05955)

$0\nu\beta\beta$  decay:

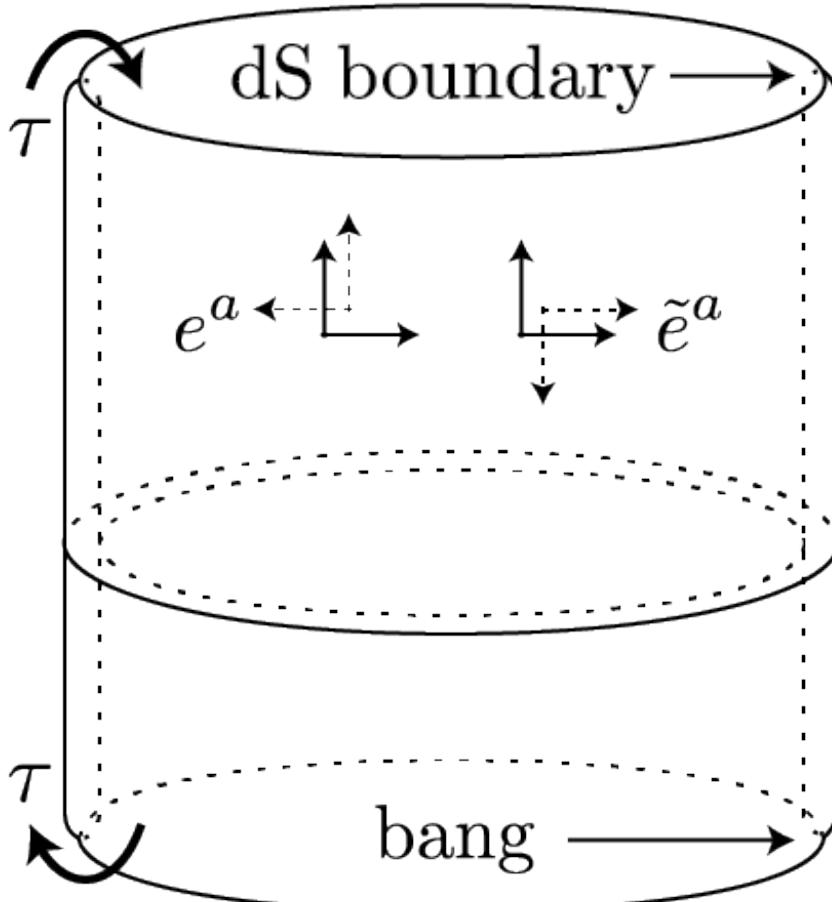


(Dell'Oro et al, arXiv:1601.07512)

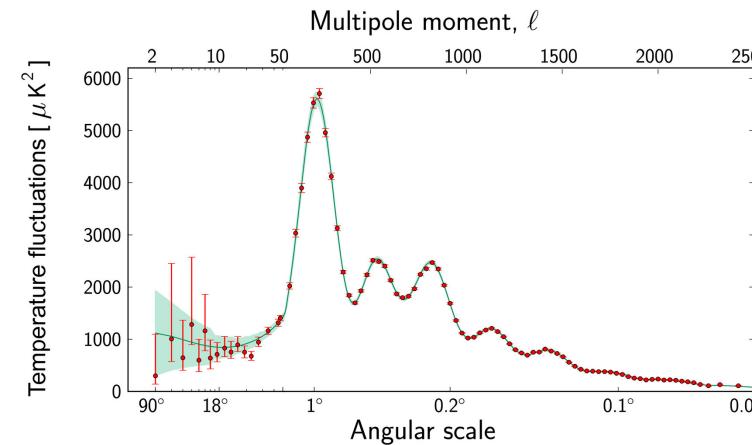
Prediction 2: dark matter is cold

Prediction 3: no primordial gravitational waves

# Consider fields on this 2-sheeted spacetime



- solns respecting background symmetry
- analytic at the bang  $\Rightarrow$  reflecting b.c.
- $\Rightarrow$  no primordial vector perturbations
- $\Rightarrow$  Neumann b.c.'s for scalar perturbation  $\zeta$
- essential singularity at dS boundary  $\Rightarrow$  no b.c.
- $\Rightarrow$  thermodynamic arrow of time



# The gravitational entropy of the universe: LB and N. Turok, arXiv:2210.01142

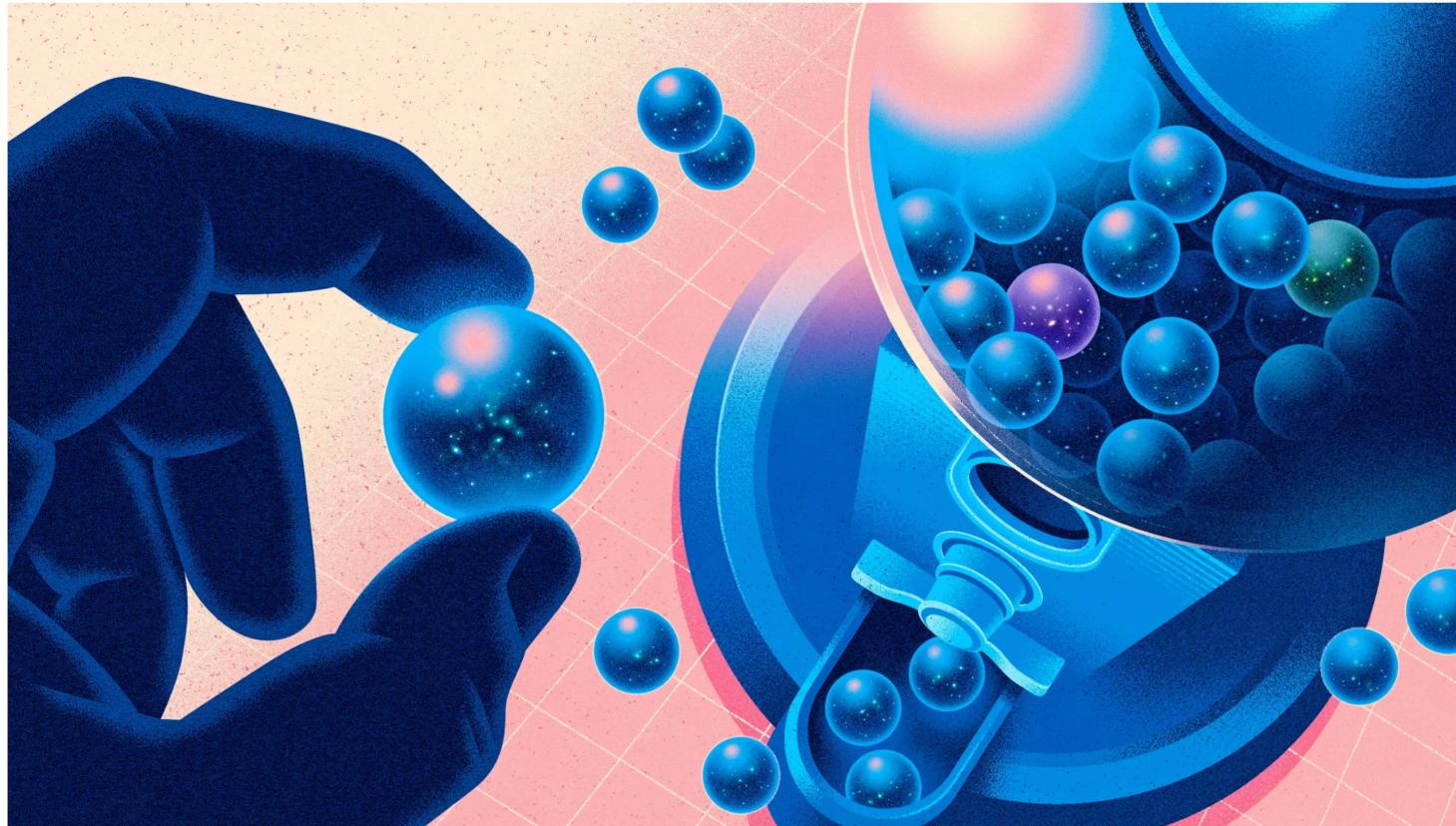
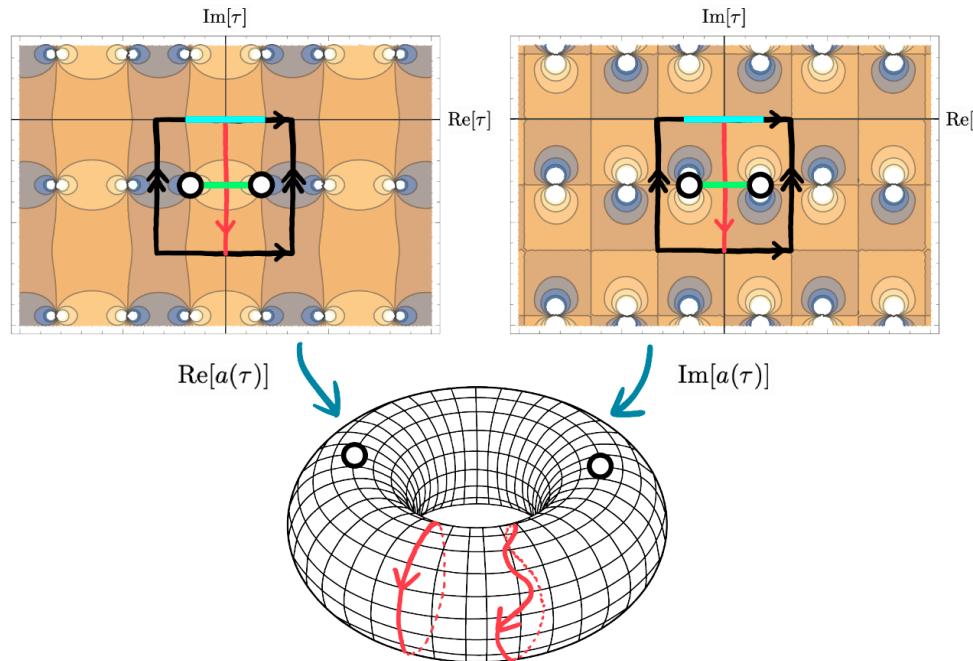


Image Credit: Quanta Magazine

# Step 1. Obtain general solution for the cosmic scale factor:

$$H^2 = \frac{8\pi G}{3} \left( \frac{r}{a^4} + \frac{\mu}{a^3} + \lambda \right) - \frac{\kappa}{a^2}$$



Step 2. Obtain general formula for the *gravitational* entropy of an FRW universe:

Flat universes and tiny positive Lambda are favoured!

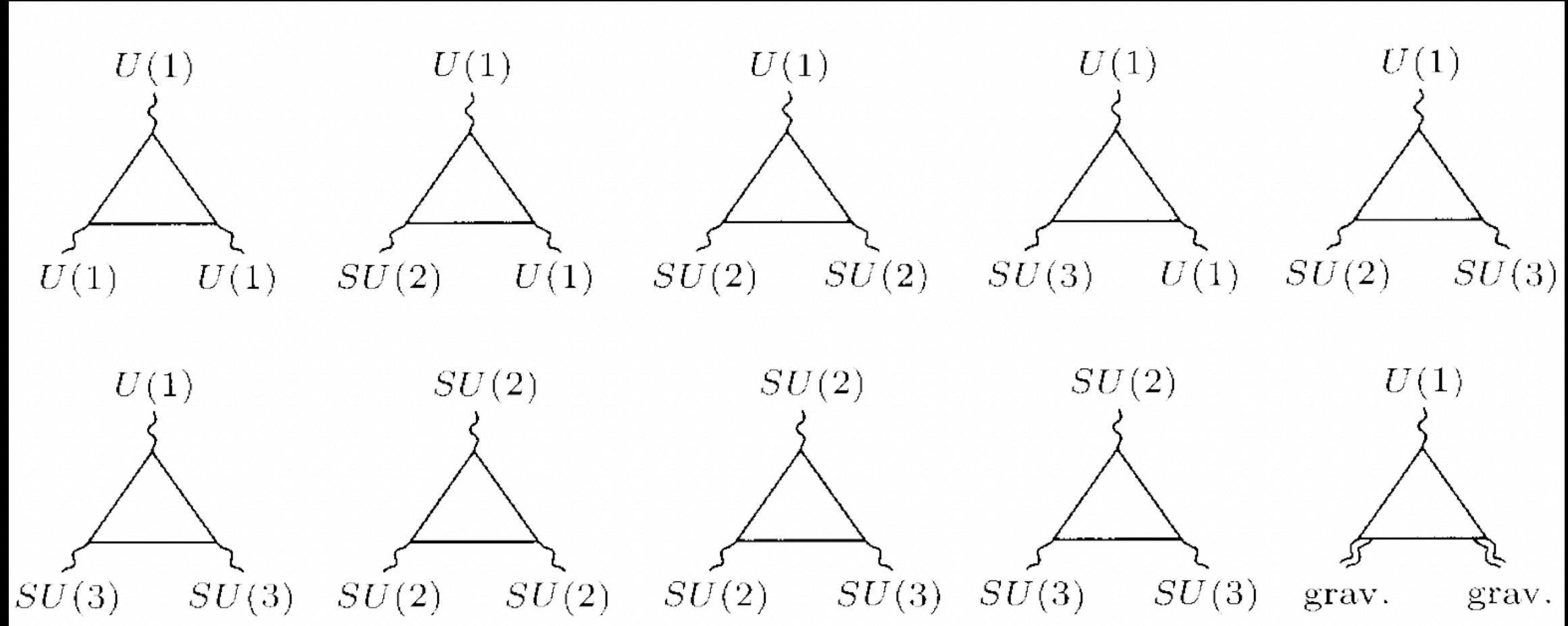
Step 3. Add cosmological perturbations (small inhomogeneities and isotropies):

If the Big Bang is a mirror, these cost entropy!

Grav. entropy is largest for universes like ours:  
homogeneous, isotropic and flat  
(with tiny positive Lambda)!

A measure on the  
space of universes.

In Standard Model, gauge and gravitational anomalies must cancel:



# What about local scale (“Weyl”) invariance?

- Emphasized by Weyl, Dirac, Dicke, ..., 't Hooft:
  - Natural generalization of diff invariance (gen. covariance)
- Ignoring Higgs, Standard Model is *\*classically\** Weyl invariant.
- But Weyl symmetry is anomalous:

$$\langle T_{\mu}^{\mu} \rangle = c C^2 - a E$$

# Cancelling Weyl anomalies and Vacuum Energy?

$$\begin{aligned} a &= \frac{1}{360(4\pi)^2} \left[ n_0 + \frac{11}{2}n_{1/2} + 62n_1 \right] \\ c &= \frac{1}{120(4\pi)^2} \left[ n_0 + 3n_{1/2} + 12n_1 \right] \end{aligned}$$

$$E_{\mathbf{k}} = \frac{\hbar\omega}{2} \left[ n_0 - 2n_{1/2} + 2n_1 \right]$$

# Cancelling Weyl anomalies and Vacuum Energy?

$$\begin{aligned} a &= \frac{1}{360(4\pi)^2} \left[ n_0 + \frac{11}{2}n_{1/2} + 62n_1 - 28n'_0 \right] \\ c &= \frac{1}{120(4\pi)^2} \left[ n_0 + 3n_{1/2} + 12n_1 - 8n'_0 \right] \end{aligned}$$

$$E_{\mathbf{k}} = \frac{\hbar\omega}{2} \left[ n_0 - 2n_{1/2} + 2n_1 + 2n'_0 \right]$$

$$S_4[\varphi] = \frac{1}{2} \int d^4x \sqrt{g} \varphi \Delta_4 \varphi \quad \Delta_4 = \square^2 + 2R^{\mu\nu} \nabla_\mu \nabla_\nu - \frac{2}{3}R \square + \frac{1}{3}(\nabla^\mu R) \nabla_\mu$$

See Fradkin and Tseytlin, Nucl. Phys. B203 (1982), 157.

# Cancelling Weyl anomalies and Vacuum Energy?

$$\begin{aligned} a &= \frac{1}{360(4\pi)^2} \left[ n_0 + \frac{11}{2}n_{1/2} + 62n_1 - 28n'_0 \right] \\ c &= \frac{1}{120(4\pi)^2} \left[ n_0 + 3n_{1/2} + 12n_1 - 8n'_0 \right] \end{aligned}$$

$$E_{\mathbf{k}} = \frac{\hbar\omega}{2} \left[ n_0 - 2n_{1/2} + 2n_1 + 2n'_0 \right]$$

$$n_{1/2} = 4n_1, \quad n'_0 = 3n_1, \quad n_0 = 0.$$

Matches standard model!

$$\begin{aligned} n_1 &= 8 + 3 + 1 = 12 \\ n_{1/2} &= 3 \times 16 = 48 \end{aligned}$$

# Dimension-zero scalars: notable features

$$\langle \varphi(t, \mathbf{x}) \varphi(t, \mathbf{x}') \rangle = \int \frac{d^3 k}{(2\pi)^3} e^{i \mathbf{k} \cdot (\mathbf{x} - \mathbf{x}')} \frac{1}{4k^3}$$

Scale invariant!

No new local degrees of freedom.

See Bogoliubov, Logunov, Oksak & Todorov (1990) and V.O. Rivelles (2003).

# Primordial Power Spectrum (arXiv:2302.00344)

$$\mathcal{P}_{\mathcal{R}}(k) = \mathcal{P}_{\mathcal{R}}(k_{pl}) (k/k_{pl})^{n_s-1}$$

$$n_s-1 \approx -\frac{7\alpha_3}{\pi} \qquad \mathcal{P}_{\mathcal{R}}(k_{pl}) = \frac{3^2 5^3}{7(2\pi)^4} \frac{1}{\mathcal{N}_{eff}^2} \left( \frac{125}{108} \alpha_Y^2 - \frac{95}{72} \alpha_2^2 - \frac{49}{6} \alpha_3^2 \right)^2$$

(SM parameters from Buttazzo et al arXiv:1307.3536])

$$n_s = 0.958 \qquad (\text{Planck : } n_s = 0.9587 \pm 0.0056)$$

$$\mathcal{P}_{\mathcal{R}}(k_*) = 12.9 \pm 4.5 \times 10^{-10} \quad (\text{Planck : } \mathcal{P}_{\mathcal{R}}(k_*) = 21 \times 10^{-10})$$

# Summary

Analytic extension of the cosmological solution of the Einstein equations leads to:

- 1) Big Bang as a mirror
- 2) formula for gravitational entropy

These ideas yield new explanations/predictions for various observed features of our universe, including:

- 1) dark matter
- 2) thermodynamic arrow of time
- 3) absence of tensor perts (gravitational wave)
- 4) absence of vector perts (vorticity)
- 5) Neumann initial conditions for scalar (density) perts
- 6) The homogeneity, isotropy and flatness (and smallness of Lambda)

Finally, we saw how (without introducing new d.o.f.) 36 dimension-zero scalars can:

- 1) cancel both Weyl anomalies and the vacuum energy
- 2) explain 3 generations
- 3) yield a scale invariant power spectrum (without inflation)

Much still to be understood!

Thank you for listening!